

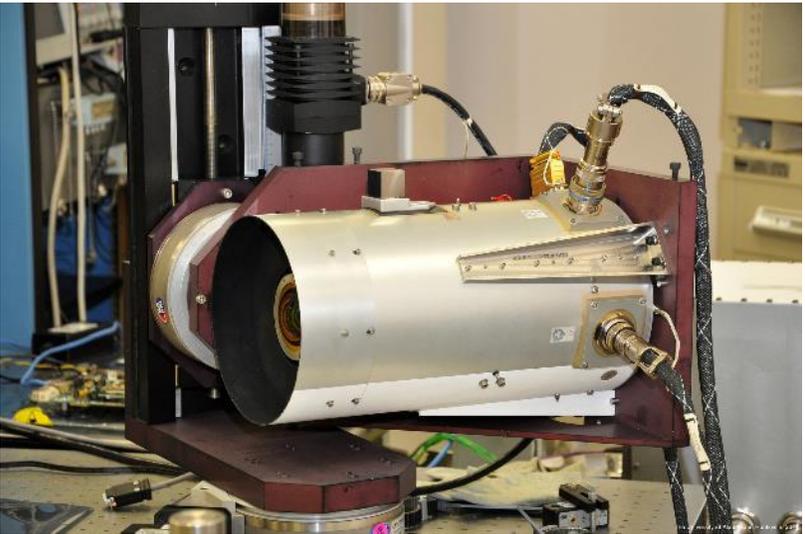
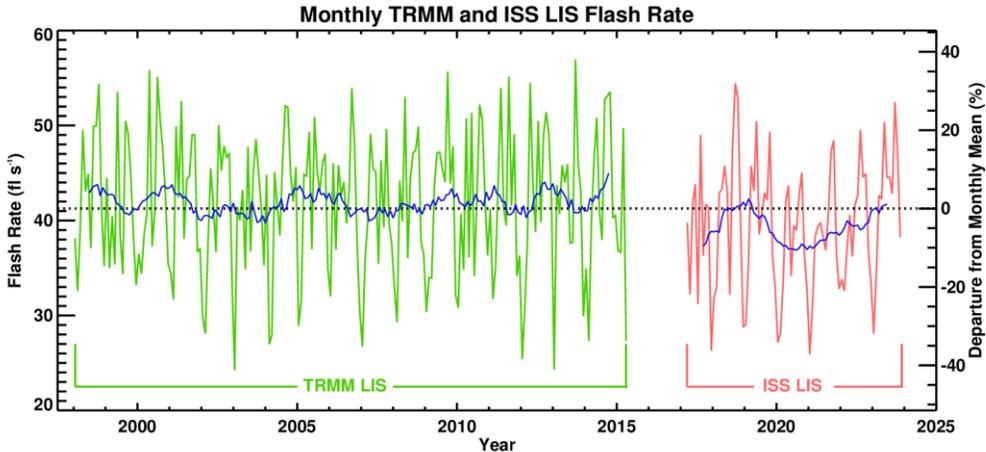
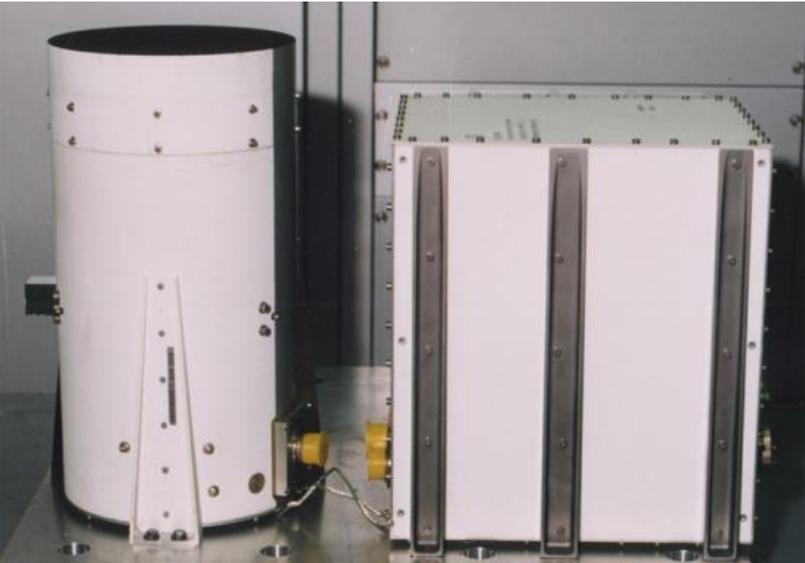


Simulations to inform NASA's Future Lightning Missions

Patrick Gatlin, M. Quick, T. Lang, J. Remington (NASA MSFC), Collaborators: S. Behnke, H. Edens (LANL), P. Bitzer (UAH), D. Mach (USRA), T. Mansell (NSSL)

*Presentation at Lightning Modeling Grand Challenge
Workshop, Albuquerque, NM, April 1-3, 2024*

Lightning Imaging Sensor (LIS) decommissioned on November 16, 2023



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3-D monitoring of lightning with optical and VHF sensors in LEO

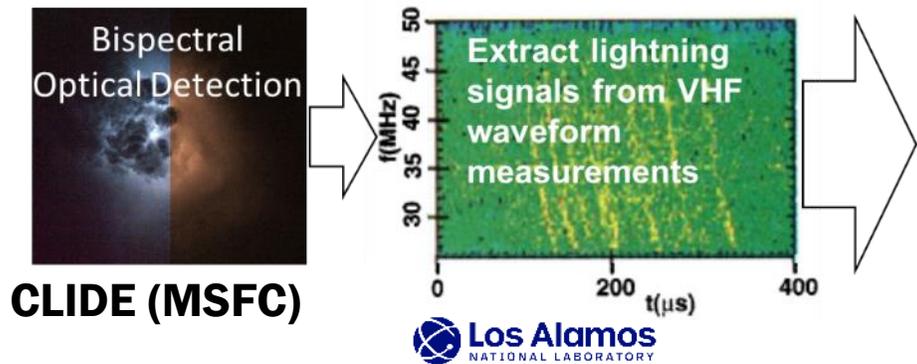
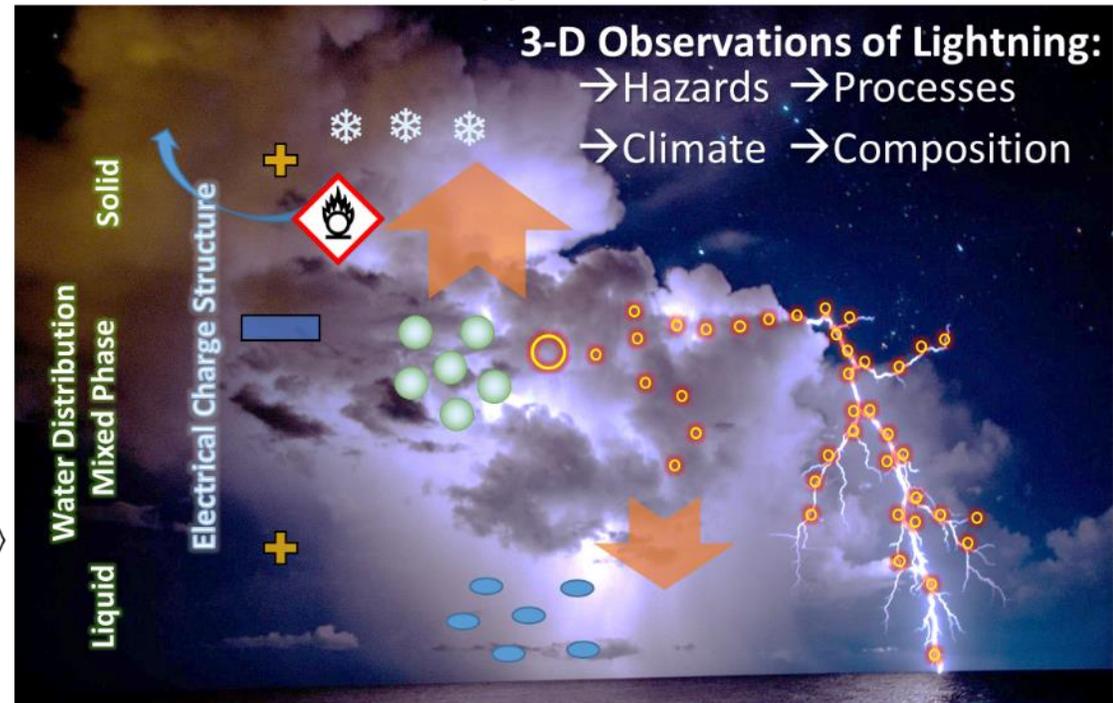
CubeSpark is a constellation of small satellites acting as a 3D lightning mapping network in space:

- *VHF radio measurements to map lightning structure inside clouds*
- *Bispectral, high-resolution optical measurements to enhance detection of lightning in severe and anomalous thunderstorms and flashes that extend upward from cloud-top*



Measurement Concept

Enabled Science and Applications



CLIDE (MSFC)

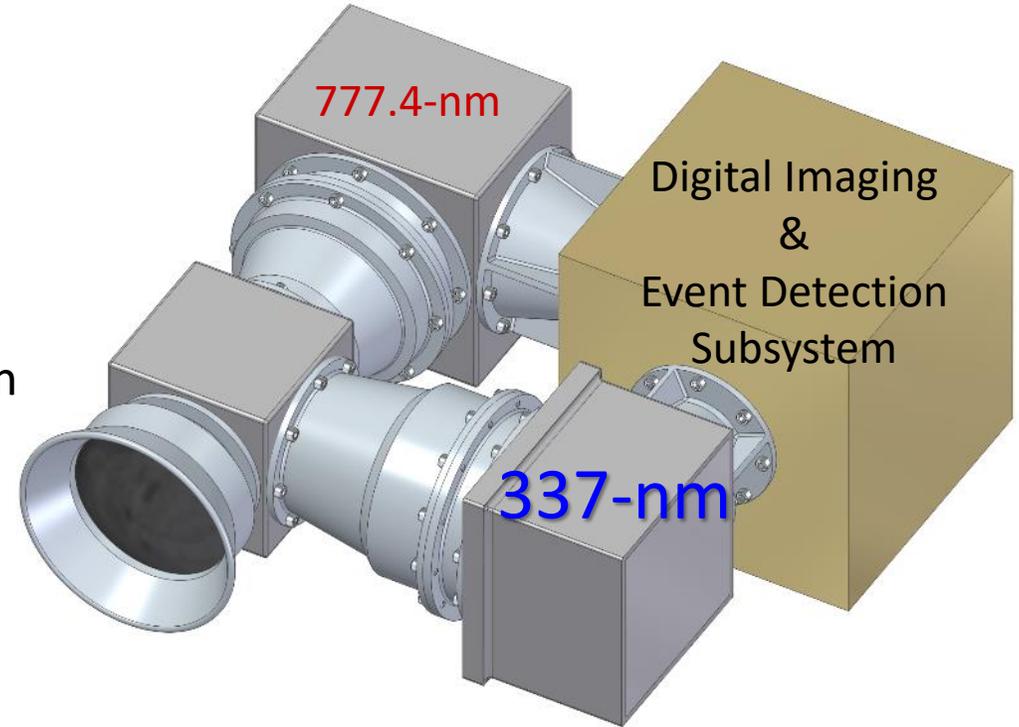


CubeSat Lightning Imaging and Detection Experiment (CLIDE)



- Objective: Improve the detection of small and optically dim lightning flashes that frequent intense thunderstorms
- Dual-wavelength:
 - 777.4-nm (OI multiplet—leaders)
 - 337-nm (N₂ SPS—streamers)
- Digital Imager and Event Detector:
 - CMOS Image Sensor developed for lightning detection
 - 432 x 420 pixel array (<2 km resolution from LEO)
 - 2000 frames per second
 - Enhance QE at 337-nm via backside processing (MBE+AR recipe)
- MSFC designing for use on small satellite missions (e.g., CubeSpark, Bushfire Monitoring)
- Funded by NASA ESTO-IIP21 as Instrument Concept Demo.
- Current TRL: 3

CLIDE Design Concept



How do we quantify observing requirements for such a mission?



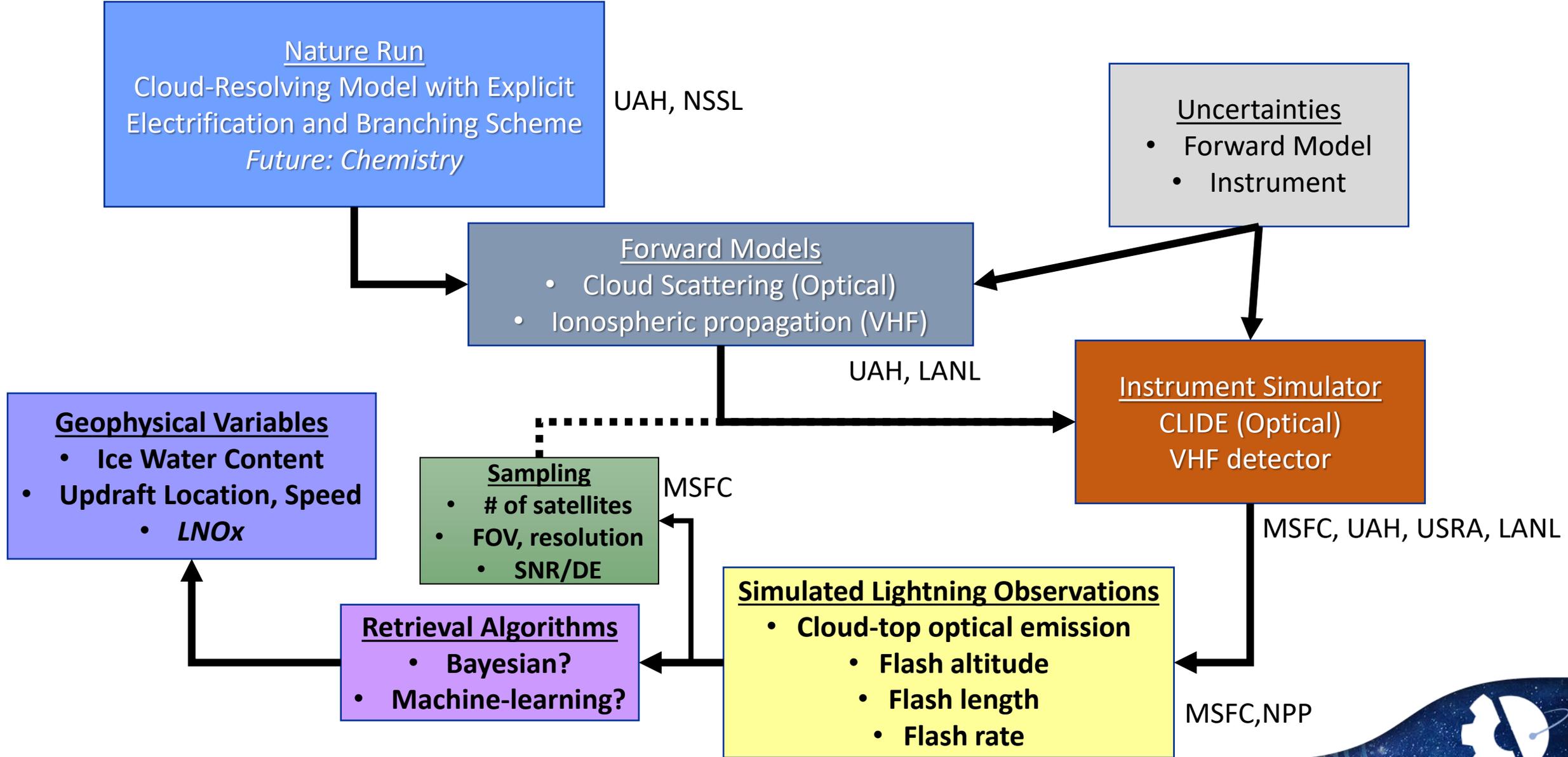
→ Observing System Simulation Experiment (OSSE)

OSSE Examples	Question Asked
Sampling	How often and with what resolution do I need to sample a feature of interest?
Retrieval	How well can the measurement estimate the geophysical variable of interest, including its uncertainties?
Process	What measurements are needed to characterize a process of interest?
Forecast	How much will the new observations improve a weather forecast?
Climate	Do these observations allow us to better constrain climate forcings or response?

Modified from CCP OSSE by D. Posselt (JPL)



Satellite-based Lightning Observing System Simulation Experiment (OSSE)



Nature Run: CRM that simulates lightning flash geometry

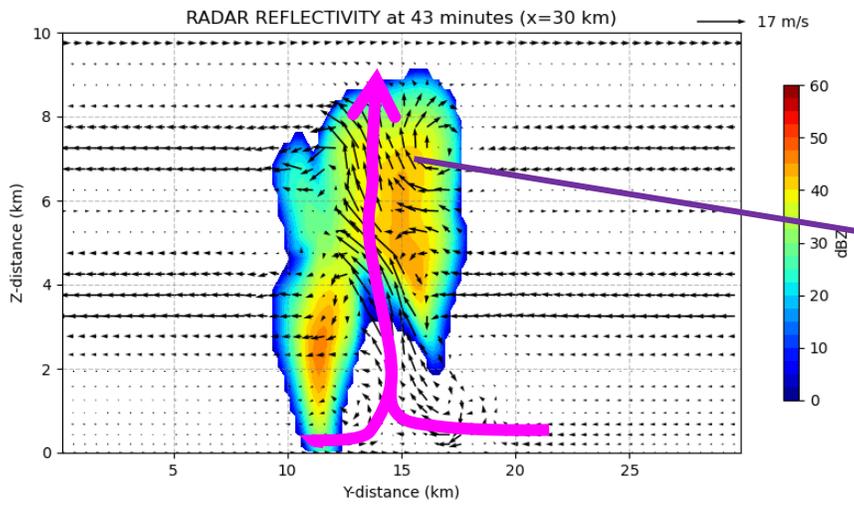
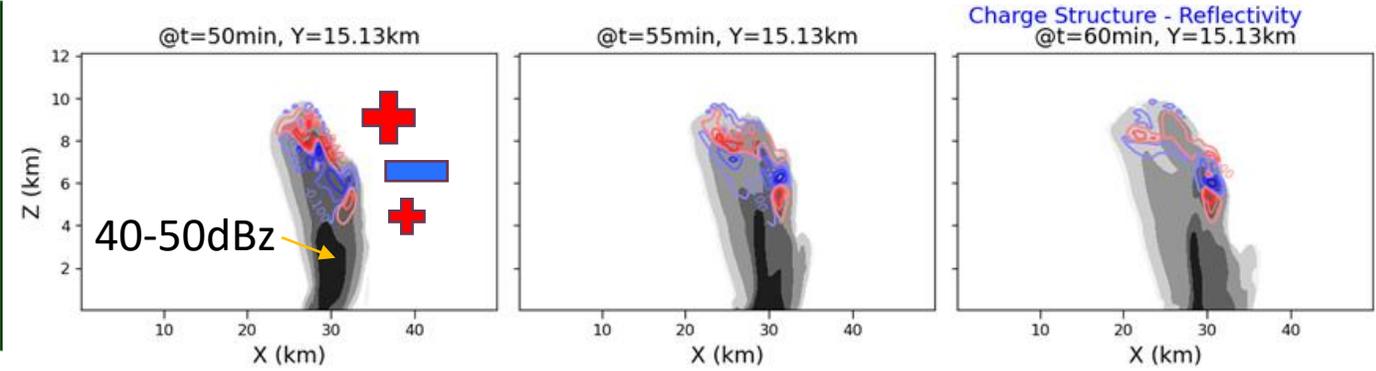


NSSL Collaborative Model for Multiscale Atmospheric Simulation (NCOMMAS; Mansell et al. 2010)

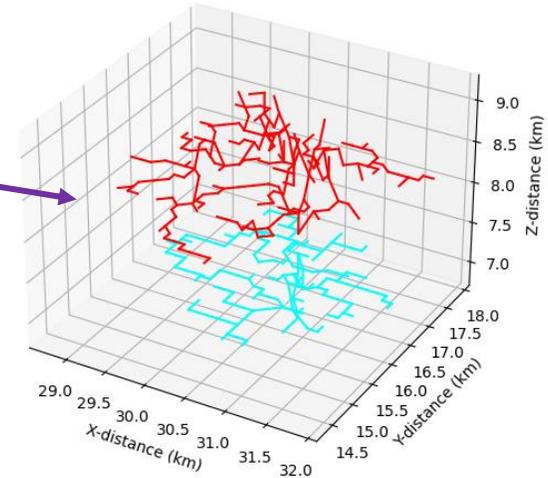
- Non-hydrostatic, convection-allowing model (Wicker & Wilhelmson 1995; Cogniglio et al. 2006)
- 2-moment microphysics (parameterizes q and n for cloud droplets, raindrops, ice crystals, aggregates, graupel; Mansell & Ziegler 2013+)
- Storm electrification scheme (parameterizes ion & hydrometeor charging; Mansell et al., 2005+)
- Discrete breakdown model (parameterizes geometry of lightning channels; Mansell et al., 2002+)

Migrating to GPU and Porting to NU-WRF

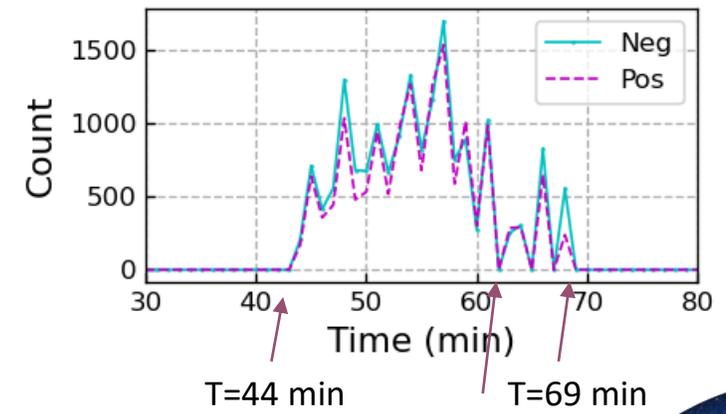
Simulation: Florida T-storm on Aug 9, 1991
Domain Size: 50 km x 30 km x 24 km
Grid Spacing: 250 m x 250 m x 100-500 m
Duration: 90 minute
Time Step: 4 sec
Output: 60 sec



Simulated Lightning Flash Structure



Simulated flash segments



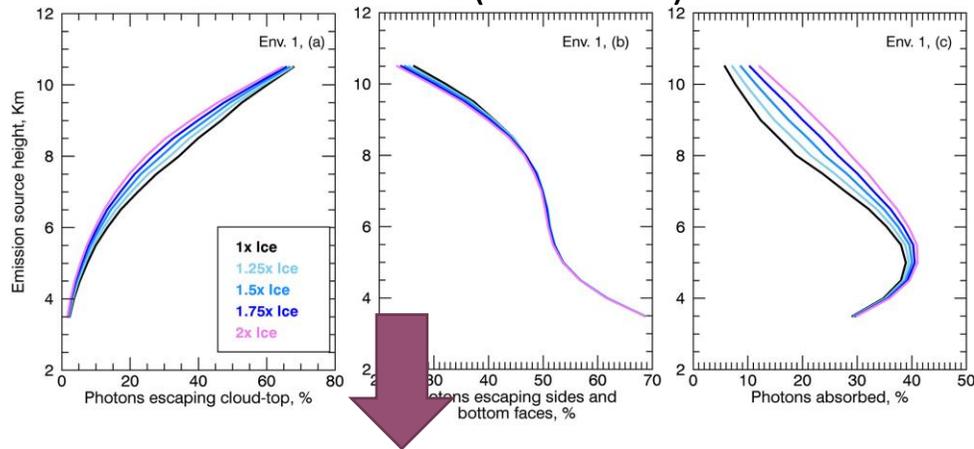
Forward Models: Cloud-top Optical Emission



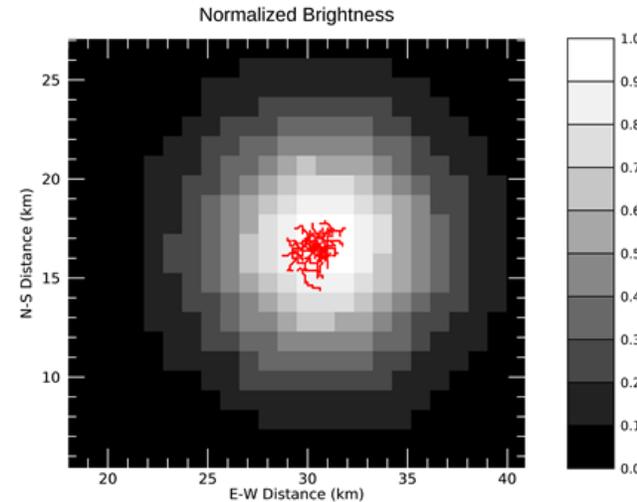
Monte Carlo Photon Pathway Model

- Simulates light emitted by lightning channel (line source)
- Lightning radiative transfer model: Thomson & Krider (1982); Koshak et al. (1994); Light et al. (2001)
- Photon extinction by inhomogeneous environment: Brunner & Bitzer (2020)

1-D Photon Extinction (1-D cloud)



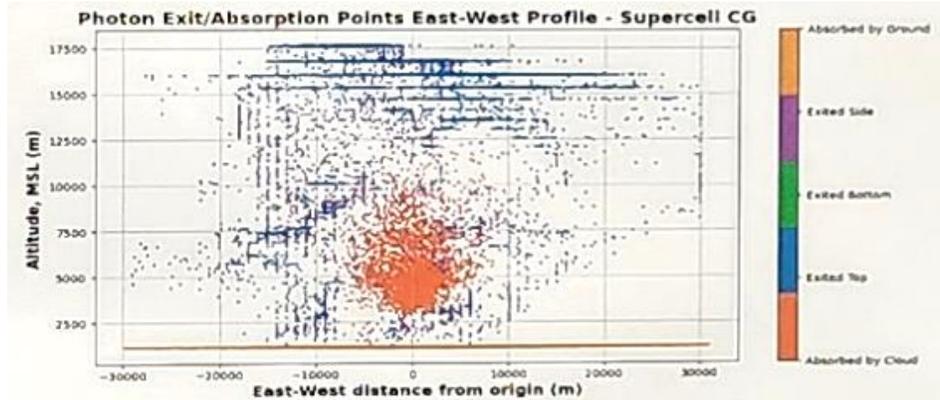
Cloud-top optical emission at 777.4-nm
Simulated lightning flash from Nature Run



Example:
 $r = 10 \mu\text{m}$
 $N_T = 10^8/\text{m}^3$
 $z_{CT} = 12\text{-km}$

J. Burchfield (UAH)

3-D Photon Extinction Model



(e.g., Wolff et al. AirForce 2023 AGU poster)

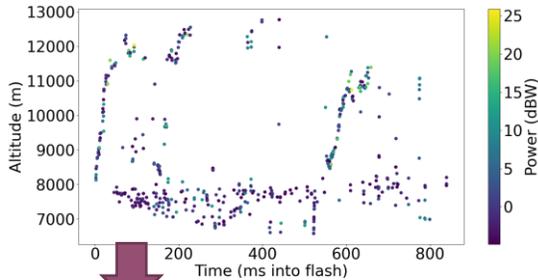


Forward Models: Ionospheric propagation of VHF signal



Los Alamos National Lab VHF signal propagation model

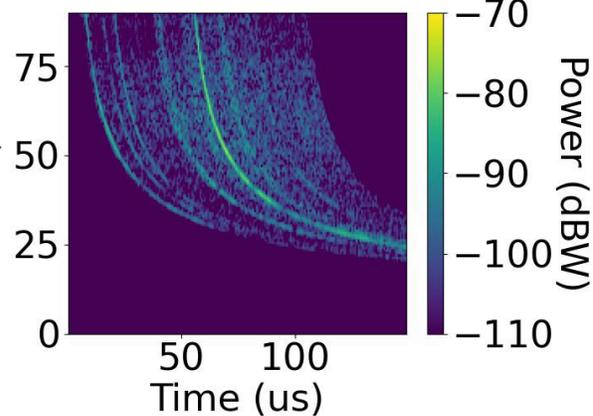
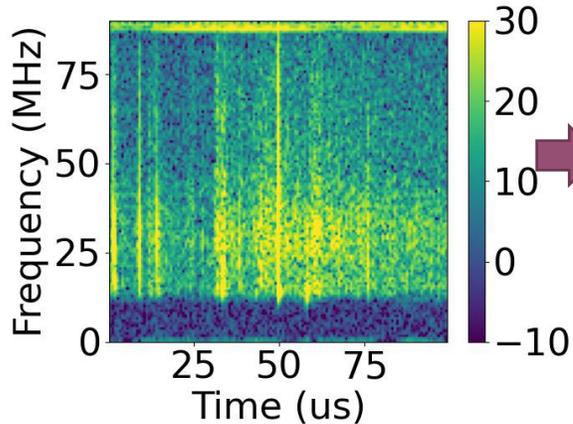
- Simulates propagation of RF wave through ionosphere
- Accounts for polarization of wave and ionospheric dispersive effects on it



LMA depiction of flash corresponding to VHF interferometric measurements used for this simulation

Pre-ionosphere impulses

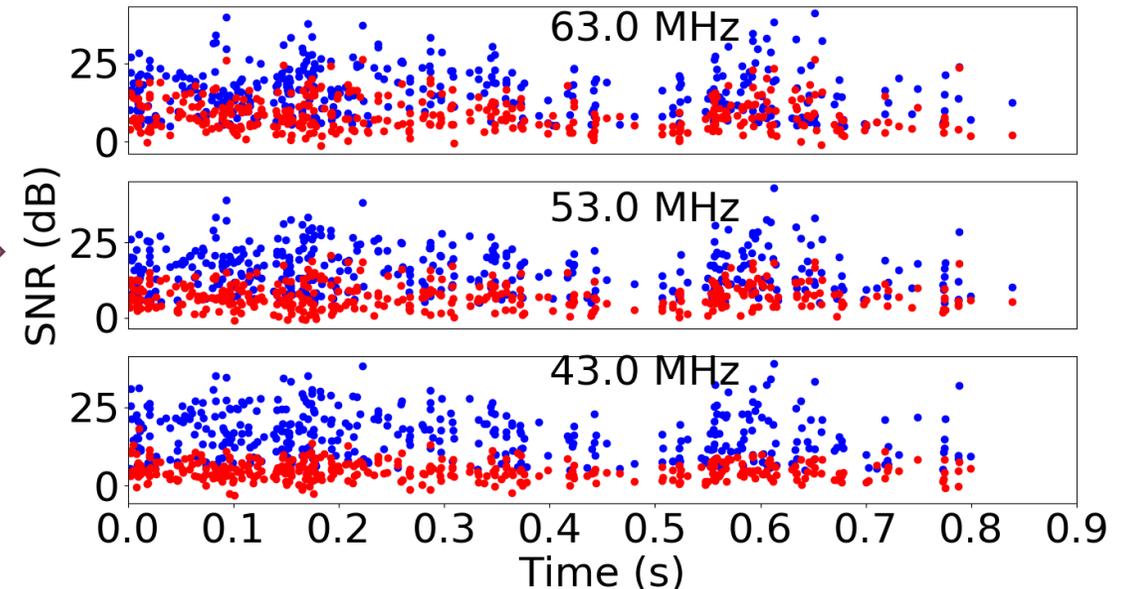
Impulses dispersed by ionosphere



Instrument Simulator: VHF Lightning Detector

- Noise estimates to compute SNR obtained from FORTE mission data (Burr et al., 2004)

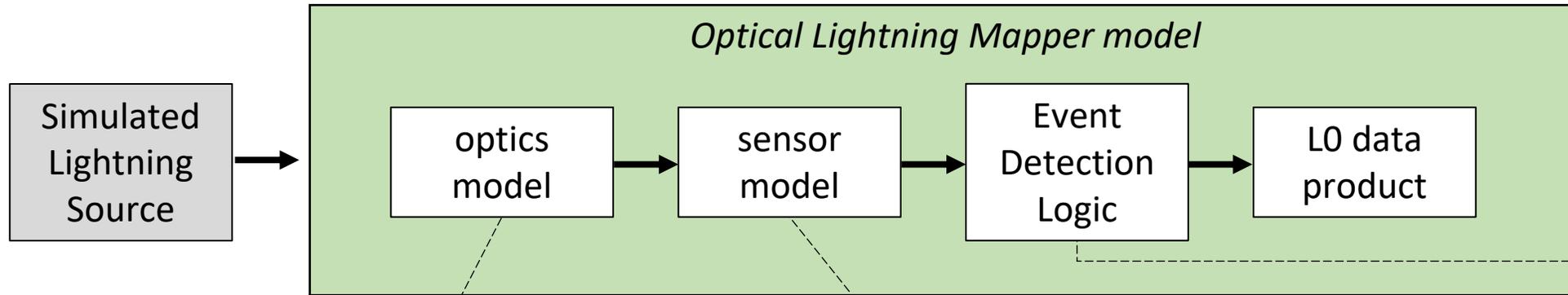
SNR at 530-km altitude (gets at detection capability)



• Original signal
• Post-ionosphere signal

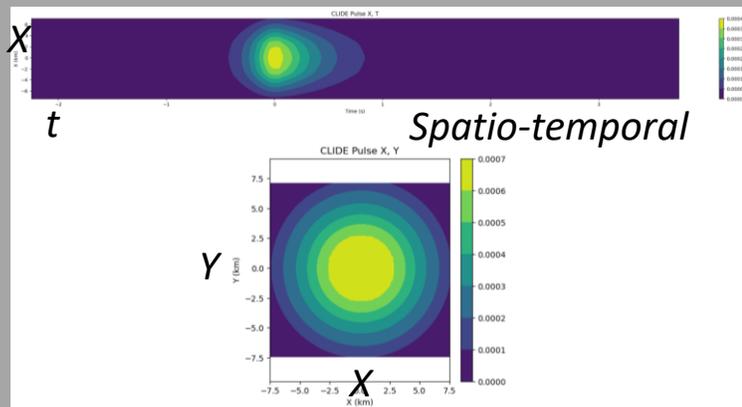
Results presented at AGU2023 by N. Pailoor et al. (LANL)





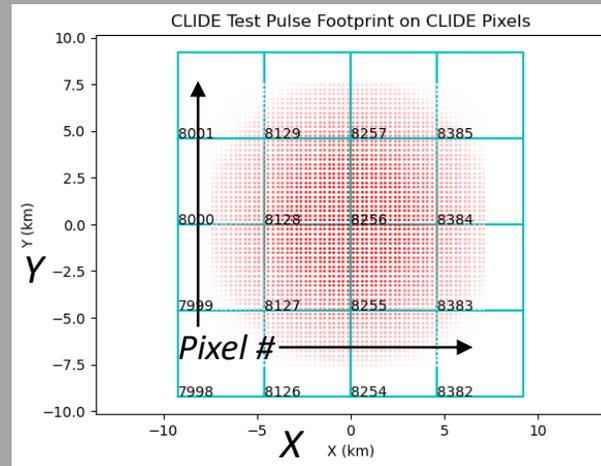
Optics Model

Simulates cloud-top optical emissions imaged at optical focal plane

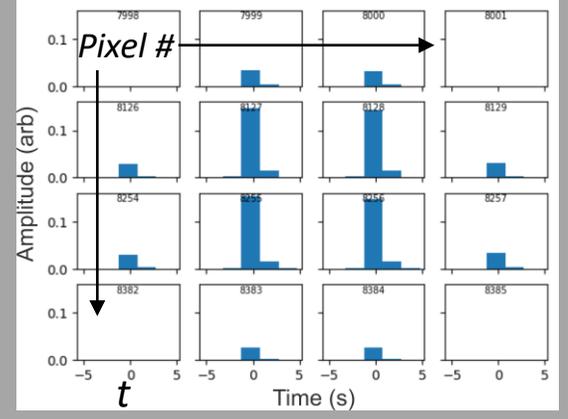


CLIDE Sensor Model

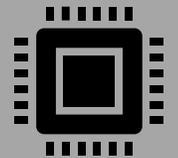
Simulates pixelated image observed by sensor



CLIDE Pixel Integrated Responses



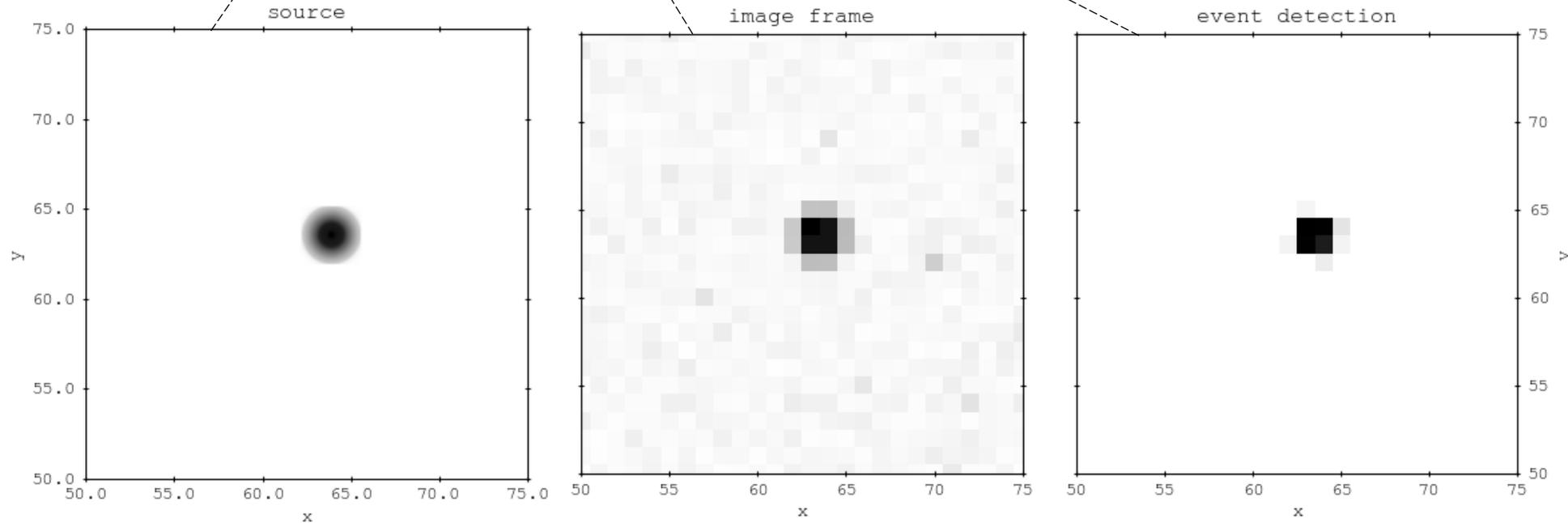
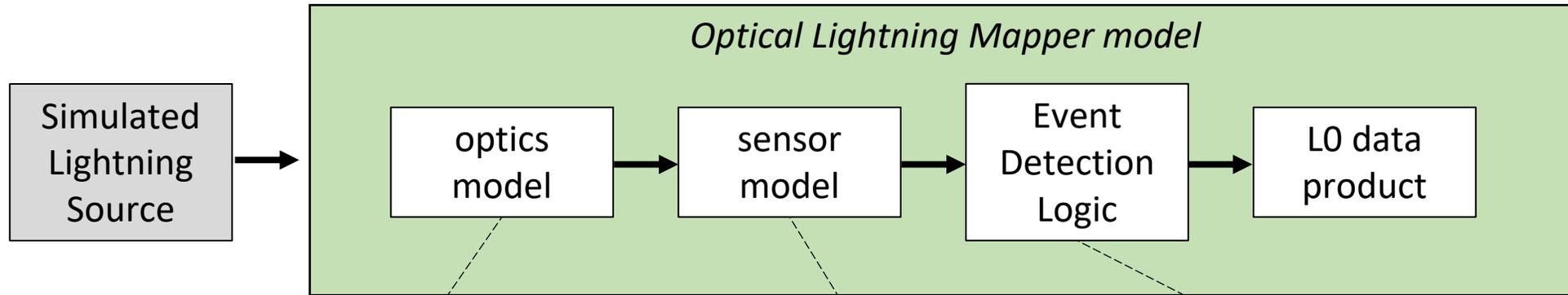
Event Detection Simulator



M. Quick (MSFC); D. Mach (USRA)



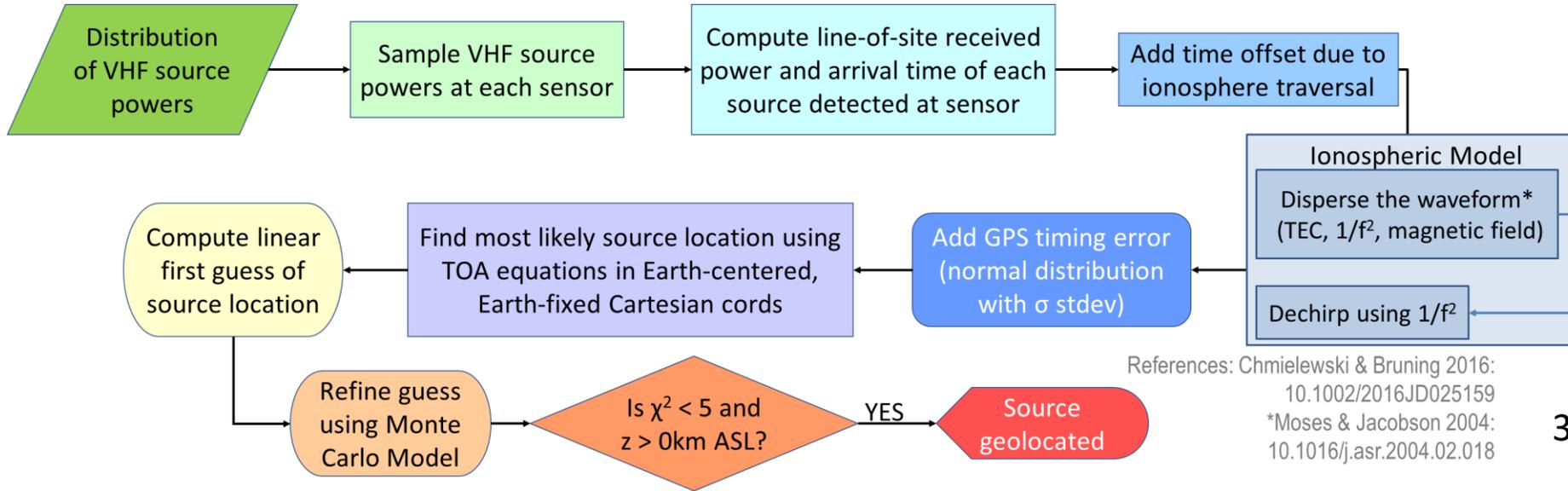
Example: Optical Lightning Event Detector



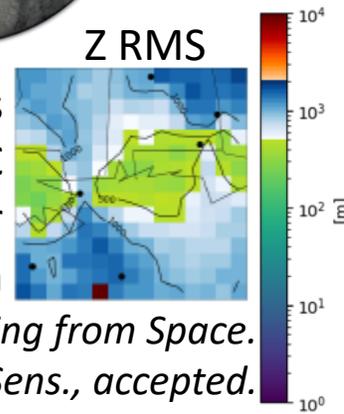
M. Quick (MSFC); D. Mach (USRA)



Simulated Observations: VHF Lightning Mapping from Space (e.g., CubeSpark)



6-satellites
30 vTEC
3-D RMS error
< 1-2 km



References: Chmielewski & Bruning 2016: 10.1002/2016JD025159
*Moses & Jacobson 2004: 10.1016/j.asr.2004.02.018

Remington et al., 2024: Simulated Feasibility of 3D Lightning Mapping from Space. IEEE Trans. Geosci. Rem. Sens., accepted.

VHF sources from lightning (e.g., LMA or Nature Run)

Simulated 3D Spaceborne Observation
Multi-satellite RF TOA Single-sat: TIPP+optical (32%)

Simulated 3D Flash Geometry

